

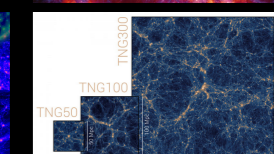
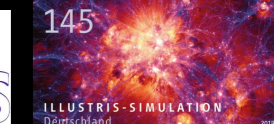
TAKING HOD MODELING TO THE NEXT LEVEL:

RESULTS FROM SDSS AND TESTS WITH HYDRODYNAMIC SIMULATIONS

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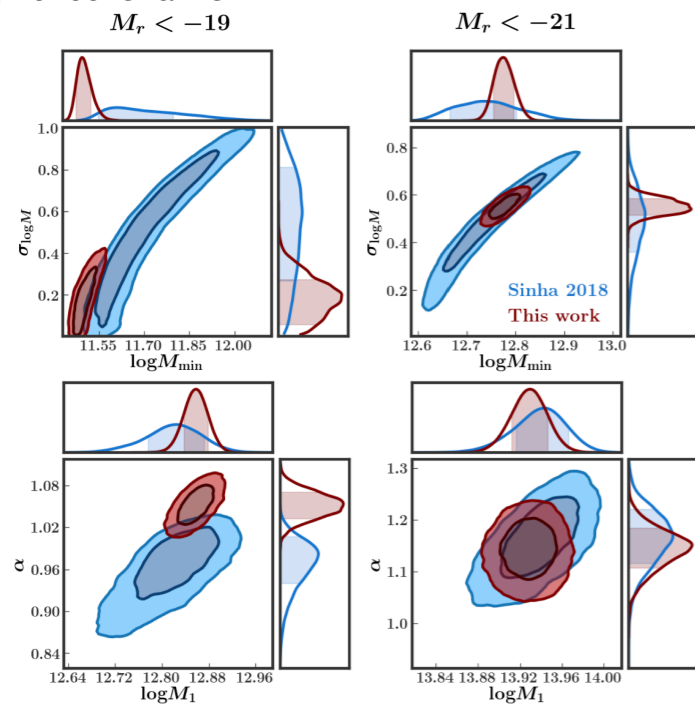
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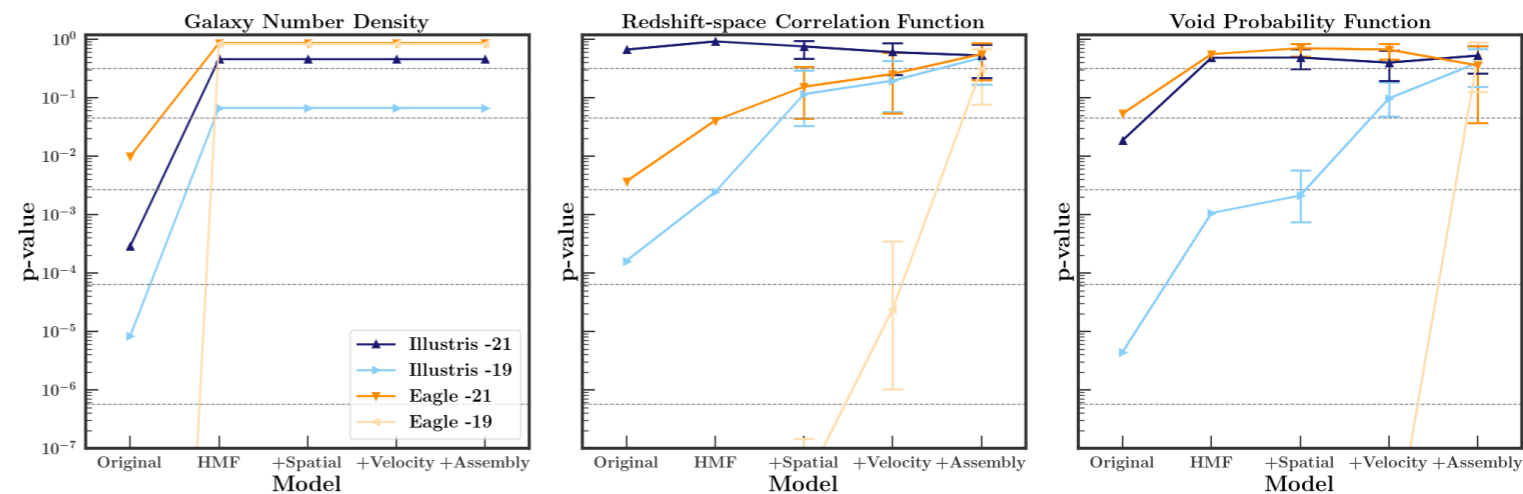
RESULTS FROM SDSS

TESTS WITH HYDRODYNAMIC SIMULATIONS

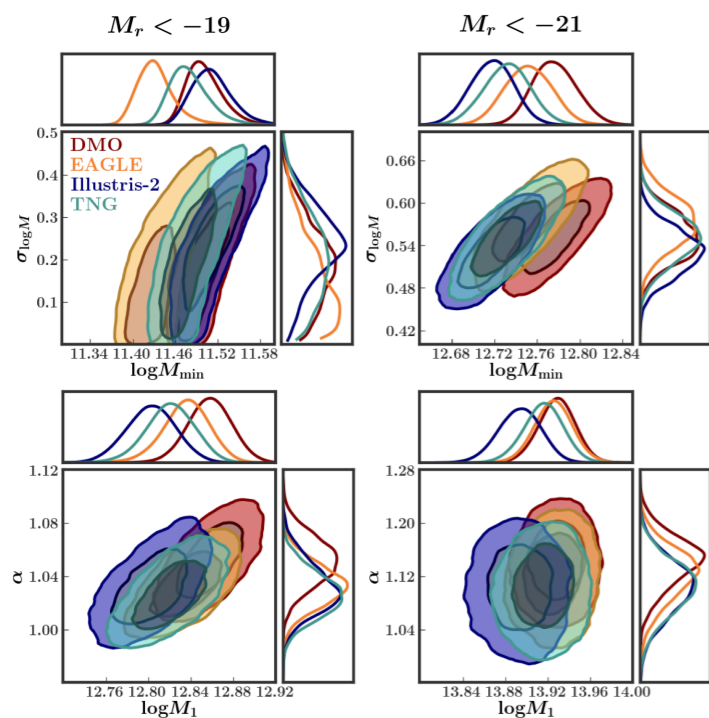
1 Szewciw, et al. (in prep): More clustering statistics leads to tighter constraints.



2 Beltz-Mohrmann, et al. (submitted to MNRAS): p-values for different clustering measurements on **Illustris** and **EAGLE** galaxies, after different modifications to the halo model and simulations. **Illustris** and **EAGLE** are very different. Faint galaxies are more affected by spatial, velocity, and assembly bias.



4 Szewciw, et al. (in prep): HOD parameter constraints change depending on how the halo masses are corrected.



3 Beltz-Mohrmann, et al. (in prep): halo mass ratios in hydro vs. DMO in **Illustris**, **EAGLE**, and **TNG**

